## Human vision and the CIE chromaticity diagram

The CIE chromaticity diagram is instrumental in colour theory. It is widely used to describe the performances of lightings or colour cameras, and more generally any tool aimed at faithfully reproducing colours as perceived by human vision. In this document, we show a practical use of this diagram through colour reproduction of the visible spectrum.

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## 1. Colour

The word "colour" is associated with the human perception of the visible light spectrum which covers the range of electromagnetic wavelengths from around 380 nm to about 740 nm . Admittedly, this is a somewhat rather abrupt approach to the notion of colours. Instead, we can tackle the subject starting from an experience that everyone is familiar with: Colour is most intuitively associated with the additive combination of the primary colours red, green and blue (RGB model). More technically, in terms that should sound for those familiar with image editing softwares, colour can be described in terms of hue, saturation, and lightness (HSL), which is an alternative representation of the RGB model. The hue value is associated with the dominant wavelength of a colour (e.g., red, orange, yellow, green, cyan, blue, indigo, violet for spectral colours) or with the complementary wavelength for non-spectral colours (purple colours obtained by the combination of red and blue primaries). Saturation provides the degree of purity of a colour, i.e. whether it appears pale or deeply saturated. Finally, lightness quantifies the intensity of a colour in terms of darkness/lightness.

All these notions may seem rather abstract without illustrations. Fortunately, they are admirably brought together in the CIE chromaticity diagram which we will discuss in this document. Before getting into the heart of the topic, we shall first briefly introduce the photopic luminosity function (section 2) - the fundamental function of photometry science - because it is needed for the construction of the CIE chromaticity diagram.

For the sake of brevity, several complex technical concepts will be covered only succinctly in this document.

## 2. The photopic luminosity function

The photopic luminosity function, $\mathrm{V}(\lambda)$, quantifies the diurnal vision of the average human eye. It was adopted in 1924 by the International Commission on Illumination (CIE, French acronym for Commission Internationale de l'Eclairage) as the standard to define the typical spectral response of the human eye to various wavelengths of light under well-lit conditions. This function is empirical since it is based on a physiological assessment. Nevertheless, it is central in photometry as it establishes the conversion from radiometric to photometric quantities. For instance, $\mathrm{V}(\lambda)$ is used to convert radiant flux $\Phi_{\mathrm{e}}$ (measured in watts, W ) into luminous flux $\Phi_{\mathrm{v}}$ (measured in lumens, Im ). $\Phi_{\mathrm{v}}$ is calculated using the following equation:

$$
\Phi_{v}=K_{c d} \int_{\lambda} \frac{\partial \Phi_{e}}{\partial \lambda} V(\lambda) d \lambda=K_{c d} \int_{\lambda=380 \mathrm{~nm}}^{\lambda=740 \mathrm{~nm}} \Phi_{e, \lambda} V(\lambda) d \lambda
$$

In this equation, $\Phi_{e, \lambda}$ is the spectral flux, or radiant flux per unit wavelength. The constant $\mathrm{K}_{\mathrm{cd}}$ is defined in the SI standard (International System of Units) as follows: The luminous efficacy of monochromatic radiation of frequency $540 \times 10^{12} \mathrm{~Hz}, \mathrm{~K}_{\mathrm{cd}}$, is $683 \mathrm{~lm} / \mathrm{W}$. This frequency corresponds to the wavelength 555.17 nm (peak of sensitivity of the eye in day vision), as


Figure 1: The $2^{\circ}$ standard observer spectral luminous efficiency function for photopic vision, CIE 1924 V( $\lambda$ ) function.

| $\boldsymbol{\lambda}(\mathbf{n m})$ | $\mathbf{V}(\boldsymbol{\lambda})$ | $\boldsymbol{\lambda}(\mathbf{n m})$ | $\mathbf{V}(\boldsymbol{\lambda})$ | $\boldsymbol{\lambda}(\mathbf{n m})$ | $\mathbf{V}(\boldsymbol{\lambda})$ | $\boldsymbol{\lambda}(\mathbf{n m})$ | $\mathbf{V}(\boldsymbol{\lambda})$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 380 | 0.00003900 | 480 | 0.13902000 | 580 | 0.87000000 | 680 | 0.01700000 |
| 385 | 0.00006400 | 485 | 0.16930000 | 585 | 0.81630000 | 685 | 0.01192000 |
| 390 | 0.00012000 | 490 | 0.20802000 | 590 | 0.75700000 | 690 | 0.00821000 |
| 395 | 0.00021700 | 495 | 0.25860000 | 595 | 0.69490000 | 695 | 0.00572300 |
| 400 | 0.00039600 | 500 | 0.32300000 | 600 | 0.63100000 | 700 | 0.00410200 |
| 405 | 0.00064000 | 505 | 0.40730000 | 605 | 0.56680000 | 705 | 0.00292900 |
| 410 | 0.00121000 | 510 | 0.50300000 | 610 | 0.50300000 | 710 | 0.00209100 |
| 415 | 0.00218000 | 515 | 0.60820000 | 615 | 0.44120000 | 715 | 0.00148400 |
| 420 | 0.00400000 | 520 | 0.71000000 | 620 | 0.38100000 | 720 | 0.00104700 |
| 425 | 0.00730000 | 525 | 0.79320000 | 625 | 0.32100000 | 725 | 0.00074000 |
| 430 | 0.01160000 | 530 | 0.86200000 | 630 | 0.26500000 | 730 | 0.00052000 |
| 435 | 0.01684000 | 535 | 0.91485010 | 635 | 0.21700000 | 735 | 0.00036110 |
| 440 | 0.02300000 | 540 | 0.95400000 | 640 | 0.17500000 | 740 | 0.00024920 |
| 445 | 0.02980000 | 545 | 0.98030000 | 645 | 0.13820000 | 745 | 0.00017190 |
| 450 | 0.03800000 | 550 | 0.99495010 | 650 | 0.10700000 | 750 | 0.00012000 |
| 455 | 0.04800000 | 555 | 1.00000000 | 655 | 0.08160000 | 755 | 0.00008480 |
| 460 | 0.06000000 | 560 | 0.99500000 | 660 | 0.06100000 | 760 | 0.00006000 |
| 465 | 0.07390000 | 565 | 0.97860000 | 665 | 0.04458000 | 765 | 0.00004240 |
| 470 | 0.09098000 | 570 | 0.95200000 | 670 | 0.03200000 | 770 | 0.00003000 |
| 475 | 0.11260000 | 575 | 0.91540000 | 675 | 0.02320000 | 775 | 0.00002120 |
| 480 | 0.13902000 | 580 | 0.87000000 | 680 | 0.01700000 | 780 | 0.00001499 |

Table 1: The $2^{\circ}$ standard observer spectral luminous efficiency function for photopic vision, CIE 1924 V( $\lambda$ ) function.
calculated from the speed of light in vacuum, $c=299792458 \mathrm{~m} / \mathrm{s}$. It is noteworthy that the integration is only carried out over the wavelengths for which $V(\lambda) \neq 0$. Since $V(\lambda)$ is given by a table of empirical values (see Figure 1 and Table 1 ), it is best to complete the integration numerically. The typical range for this calculation is $380 \mathrm{~nm} \leq \lambda \leq 740 \mathrm{~nm}$.
This short incursion into photometry highlights the fundamental role played by $\mathrm{V}(\lambda)$. We will not discuss photometry further as it falls outside the scope of this technical note. The interested readers can find details on this topic in the Technical Guide "A general guide to high-efficiency LED exposure systems" released by idonus in 2019 (DOI: 10.13140/RG.2.2.29958.32323).
We also temporarily leave the photopic function $V(\lambda)$, to which we will return in section 5 .

## 3. Wright and Guild's colour matching experiments

In the late 1920's, William David Wright (Imperial College, London) and John Guild (National Physical Laboratory, Teddington) conducted a series of experiments aimed at consolidating the emerging theory of colorimetry. Both lived in the UK, and the studies they carried out were similar in many aspects, but they completed their work independently from each other. This is an important point because it was the concordance between their respective findings that motivated the CIE to define a standardised colour space based on their results.

In short, observers were asked to match monochromatic lights by mixing three primary-colour lights. Using the data collected from only 17 observers ( 10 for Wright and 7 for Guild), Wright and Guild could quantitatively measure how the monochromatic colours of the spectrum can be reproduced when beams of red, green, and blue light are added. The combination of their data, gathered by Guild in 1931, forms the basis of the CIE colour matching functions that are still in use in today's colorimetry science. Below, we describe the basic parts of Guild's experimental setup. The main difference from Wright's experiments lied in the wavelengths chosen for the primaries.

- Monochromator - The source of light (reference white) was a 40-watt automobile headlight lamp, the filament of which was focused by means of an achromatic lens and a prism on the slit of a constant deviation monochromator.
- Primaries - The 3 primaries were obtained by passing the light from an opal-bulb gas-filled lamp through red, green, and blue gelatine filters. The wavelength chosen for these primaries were 700 nm (red), 546.1 nm (green), and 435.8 nm (blue). The choice of very specific wavelengths for green and blue lied in the availability of filters specially designed to isolate these spectral lines prominent in mercury. As for the red stimulus, because the hue was found to be constant for wavelengths higher than about 670 nm , a convenient procedure was to use a very deep red filter.
- Trichromatic colorimeter - The observer was facing a square field of view divided into two rectangular portions by a horizontal line. The size of the aperture was such that the square subtended an angle of approximately $2^{\circ}$ at the observer's eye (hence, the name "2-degree standard observer"). The test field contained the monochromatic colour to be matched. The other, the matching field, contained a blend of the primaries which could be adjusted by means of three control handles, their amounts being measured on graduations attached to the shutters. If a test colour couldn't be matched, the observer could instead introduce one of the primaries into the test field (i.e., subtracted).

With this system, any monochromatic colour ( $\mathrm{C}_{\lambda}$ ) could be uniquely represented by a colour equation of the form:

$$
C_{\lambda}=r R+g G+b B, \quad r+g+b=1
$$

where $r, g$, and $b$ are the trichromatic coefficients, positive or negative, representing certain proportions of the red (R), green (G), and blue (B) primaries, respectively ${ }^{1}$. In Figure 2, we show a graph of the trichromatic coefficients plotted from the values adopted by the CIE Committee on Colorimetry in 1931. Since these coefficients are normalized, the CIE $r g$ chromaticity space can also be represented in the form of the $r g$ chromaticity diagram shown in Figure 3, the $b$ coordinate being directly deduced from the $r$ and $g$ coordinates. The curved line is the spectral locus. The red ( $1,0,0$ ), green $(0,1,0)$, and blue $(0,0,0)$ points on that curve correspond to the 3 primary colours. The white point $(\mathrm{E})$ is the Equal Energy point where $r=g=b$. Notice that the spectral locus is completely enclosed in a triangle (dashed lines) formed by the virtual points $\mathrm{C}_{\mathrm{r}}, \mathrm{C}_{\mathrm{g}}$ and $\mathrm{C}_{\mathrm{b}}$. The Equal Energy point and the vertices of the dashed triangle are located at the following coordinates:

|  | E | $\mathrm{C}_{\mathrm{r}}$ | $\mathrm{C}_{\mathrm{g}}$ | $\mathrm{C}_{\mathrm{b}}$ |
| :---: | :---: | :---: | :---: | :---: |
| $\mathbf{r}$ | $1 / 3$ | 1.27496 | -1.73930 | -0.74310 |
| g | $1 / 3$ | -0.27777 | 2.76726 | 0.14091 |
| b | $1 / 3$ | 0.00281 | -0.02795 | 1.60219 |

Table 2: Coordinates of the equal energy point and three primary colours in the CIE RGB colour space.

To draw the dashed triangle in Figure 3, the vertices were chosen arbitrarily but conscientiously. The line that passes through $C_{b}$ and $C_{r}$ is known as the alychne, a word coined by Erwin Schrödinger, meaning "no light" in Ancient Greek: it is

[^0]

Figure 2: Trichromatic coefficients $r, g$, and $b$ of the CIE 1931 standard observer.


Figure 3: CIE 1931 rg chromaticity diagram.
the line of zero luminance in the colour space. The line that passes through $\mathrm{C}_{\mathrm{g}}$ and $\mathrm{C}_{\mathrm{r}}$ is located just above the green and red primaries. As for the line that passes through $\mathrm{C}_{\mathrm{g}}$ and $\mathrm{C}_{\mathrm{b}}$, it was chosen so as to keep the size of the triangle small. On the spectral locus, the closest point is approximately that of the 504 nm wavelength.


Figure 4: CIE 1931 xy chromaticity diagram. The triangle defined by the red, green, and blue points is the gamut of the original $R G B$ space.

## 4. CIE XYZ colour space

The closed area formed by the spectral locus curve and the alychne line encompasses all the colours visible for the average human eye. In Figure 3, the right-angled triangle formed by the red, green, and blue points only contains the colours that can be synthesized with this particular choice of trichromatic primaries (i.e., $435.8 \mathrm{~nm}, 546.1$, and 700 nm colour stimuli). To get free from a particular choice of the primary colours, the CIE opted for a virtual XYZ colour space built from the virtual primaries $C_{r}, C_{g}$ and $C_{b}$. These primaries have the following coordinates in the CIE XYZ colour space: $C_{r}(1,0,0), C_{g}(0,1,0), C_{b}$ $(0,0,1)$. For anyone familiar with linear algebra, an immediate matrix transformation that comes to mind is that obtained from the trichromatic coordinates given in Table 1:

$$
M_{0}=\left[\begin{array}{ccc}
1.27496 & -1.73930 & -0.74310 \\
-0.27777 & 2.76726 & 0.14091 \\
0.00281 & -0.02795 & 1.60219
\end{array}\right]^{-1}=\left[\begin{array}{lll}
0.90878 & 0.57494 & 0.37093 \\
0.09122 & 0.41876 & 0.00548 \\
0.00000 & 0.00630 & 0.62359
\end{array}\right]
$$

However, this transformation would translate the Equal Energy point to the following xyz coordinates:

$$
E_{0}=\left[\begin{array}{c}
E_{0, x} \\
E_{0, y} \\
E_{0, z}
\end{array}\right]\left[\begin{array}{c}
0.61822 \\
0.17182 \\
0.20996
\end{array}\right]
$$

The CIE wanted the coordinate of the Equal Energy point to remain unchanged through the transformation RGB $\rightarrow$ XYZ. The xyz coordinates being normalized, one can see that the appropriate matrix transformation, $M$, is of the form:

$$
\begin{gathered}
M=M_{R G B \rightarrow X Y Z}=k \cdot \frac{1}{3}\left[\begin{array}{ccc}
1 / E_{0, x} & 0 & 0 \\
0 & 1 / E_{0, y} & 0 \\
0 & 0 & 1 / E_{0, z}
\end{array}\right] M_{0}=k\left[\begin{array}{llll}
0.49000 & 0.31000 & 0.20000 \\
0.17697 & 0.81240 & 0.01063 \\
0.00000 & 0.01000 & 0.99000
\end{array}\right] \\
M=\left[\begin{array}{ccc}
2.768892 & 1.751748 & 1.130160 \\
1 & 4.5907 & 0.0601 \\
0 & 0.056508 & 5.59429
\end{array}\right]
\end{gathered}
$$

where $k$ is a scaling factor that doesn't affect the $x y z$ coordinates ( $k=5.6508$, see section 5 for explanations). In this equation, the $1 / 3$ factor is a simple rewriting of the coordinates of the point of Equal Energy. The result of this transformation is shown in Figure 4. In this CIE XYZ chromaticity diagram, we have added colours to represent all the colours that can be seen by the human eye. As stated above, only the colours that are inside the triangle formed by the red, green, and blue can be synthesized with this choice of trichromatic primaries. This triangle represents the gamut of this colour space: the subset of colours which can be accurately represented with these primaries. Since most monitors and printers are designed to fulfil the standardized sRGB colour space, which has a smaller gamut than the CIE RGB colour space, part of the colours inside the triangle cannot be reproduced accurately. This remark is even more true for the colours outside the gamut triangle: they can simply not be reproduced. For this reason, the blue-cyan-green colours shown in the upper part of the graph are simply an extension of the colours calculated on the edge of the triangle.

## 5. CIE colour matching functions

So far, we have only dealt with normalized coefficients ( $r g b$ and $x y z$ ) that provide the chromaticity coordinates in a colour space. In its effort to standardize the XYZ colour space, the CIE introduced absolute colour matching functions from which all the normalized coefficients are calculated. They are written with an overbar: $\bar{x}(\lambda), \bar{y}(\lambda), \bar{z}(\lambda)$ for the $X Y Z$ space and $\bar{r}(\lambda), \bar{g}(\lambda)$, $\overline{\mathrm{B}}(\lambda)$ for the RGB space. Owing to the fundamental role played by the photopic function, the CIE imposed that $\bar{y}(\lambda)$ be the photopic function $V(\lambda)$ introduced in section 2:

$$
\bar{y}(\lambda)=V(\lambda)
$$

It can be straightforwardly shown that the other colour matching function must satisfy the following relations:

$$
\bar{x}(\lambda)=\frac{x}{y} V(\lambda), \quad \bar{z}(\lambda)=\frac{V(\lambda)}{y}(1-x-y)
$$

As for the $\bar{r}(\lambda), \bar{g}(\lambda)$, and $\bar{b}(\lambda)$ colour matching functions, because they have a physical significance, an appropriate scaling factor needed to be used for $X Y Z \rightarrow R G B$ transformation. Indeed, the $\bar{r}(\lambda), \bar{g}(\lambda)$, and $\overline{\mathrm{b}}(\lambda)$ functions were normalized so that:

$$
\overline{\mathrm{r}}(\lambda)+4.5907 \bar{g}(\lambda)+0.0601 \bar{b}(\lambda)=\mathrm{V}(\lambda)
$$

The ratios $1: 4.5907: 0.0601$ correspond to the relative luminances of the trichromatic stimuli. Hence, the parameter $k=5.6508$ that was used above without further explanation $(1+4.5907+0.0601=5.6508)$.

The colour matching functions $\bar{r}(\lambda), \bar{g}(\lambda), \bar{b}(\lambda)$ can be obtained through the matrix transformation:

$$
M_{X Y Z \rightarrow R G B}=M^{-1}=\left[\begin{array}{ccc}
2.768892 & 1.751748 & 1.130160 \\
1 & 4.5907 & 0.0601 \\
0 & 0.056508 & 5.594292
\end{array}\right]^{-1} \approx\left[\begin{array}{ccc}
0.418456 & -0.158657 & -0.082833 \\
-0.091167 & 0.252426 & 0.015707 \\
0.000921 & -0.002550 & 0.178595
\end{array}\right]
$$

The colour matching functions for the XYZ and RGB spaces are shown in Figure 5 and Figure 6, respectively.


Figure 5: CIE XYZ colour matching functions.


Figure 6: CIE RGB colour matching functions

## 6. sRGB colour space and Planckian locus

We have mentioned above the sRGB colour space which is widely used for consumer electronic devices (displays, digital cameras, printers), and which is also the standard for the Internet. Conversion from CIE XYZ to sRGB requires a scaling so that the white point is D65. The illuminant D65 is a CIE standard that is intended to represent average daylight. It was calculated from the average spectral power distribution of the Sun and has a correlated colour temperature of approximately 6500 K .


Figure 7: CIE 1931 chromaticity diagram and sRGB colour gamut. The curve passing through the white point D65 is the Planckian locus. The points highlighted on the spectral locus correspond to monochromatic lights emitted at the indicated wavelengths (in nm).

The coordinates of D65 in the CIE XYZ colour space can be approximated using the theoretical spectral radiant exitance of a black body heated at $\mathrm{T}=6503.15 \mathrm{~K}$ :

$$
M_{e, \lambda}=\frac{2 \pi h c^{2}}{\lambda^{5}}\left(e^{\frac{h c}{k T}}-1\right)^{-1}
$$

where $h$ is the Planck constant, $c$ is the speed of light in vacuum, $k$ is the Boltzmann constant, $T$ is the temperature in Kelvin, $\lambda$ is the wavelength. In practice, this equation is used to draw the Planckian locus curve. In Figure 7, we have drawn the CIE 1931 XYZ chromaticity diagram and included the Planckian locus curve. The dashed triangle shows the gamut of the sRGB colour space. The colours were calculated using the $X Y Z \rightarrow s R G B$ transformation matrix from the official sRGB documentation:

$$
M_{X Y Z \rightarrow S R G B}=\left[\begin{array}{ccc}
3.2406255 & -1.537208 & -0.4986286 \\
-0.9689307 & 1.8757561 & 0.0415175 \\
0.0557101 & -0.2040211 & 1.0569959
\end{array}\right]
$$

The GNU Octave script prepared to generate the CIE diagram of Figure 7 is provided at the end of this technical note (3 pages, grey background) ${ }^{2}$.

[^1]

Figure 8: Calculating saturated colour coordinates for sRGB colour space. The line that joins point D65 and a specific point on the spectral locus (e.g., 570 nm monochromatic wavelength) crosses the triangle gamut at a specific coordinate (e.g., $x=0.4166 ; y=0.5074$ ).

## 7. Rendering of the colour spectrum in the sRGB colour space

In Figure 7, the points on the spectral locus curve represent pure monochromatic colours visible to the average human eye. However, only the points inside the dashed triangle can be represented in the sRGB colour space. Taking this limitation into account, the best way to perform optimum rendering of the colour spectrum is to use the corresponding saturated values that are found on the edges of the gamut triangle. An example is shown in Figure 8 for 570 nm monochromatic colour (yellow). The calculation must be made using the reference white colour (for SRGB colour space, the white point is D65). The saturated colour is found straightforwardly using the line that passes through the white point and a given point on the spectral locus: It crosses the gamut triangle at a well-defined coordinate near the spectral locus. These calculated coordinates are plotted in Figure 9. Using the XYZ $\rightarrow$ sRGB transformation matrix introduced in section 6 , one can easily obtain the curve plotted in Figure 10. Notice that the procedure followed is identical to the one described in section $5: \bar{x}(\lambda), \bar{y}(\lambda), \bar{z}(\lambda)$ must be calculated prior to completing the transformation.

These linear RGB values are not the final result. Gamma correction must be applied to obtain the final non-linear sRGB values. The following formula transforms the linear values into $\operatorname{sRGB}$ :

$$
C_{s R G B}=\left\{\begin{array}{cc}
12.92 C_{\text {linear }}, & C_{\text {linear }} \leq 0.0031308 \\
1.055 C_{\text {linear }}^{1 / 2.4}-0.055, & C_{\text {linear }} \leq 0.0031308
\end{array}\right.
$$



Figure 9: Saturated colour coordinates in sRGB colour space and corresponding monochromatic wavelengths. For example, for 570 nm (yellow), the corresponding coordinates are: $x=0.4166 ; y=0.5074$; $z=0.076$.


Figure 10: linear RGB values and corresponding wavelengths for colour coordinates shown in Figure 8.
where $C$ is either $R, G$, or $B$. These gamma-corrected values are usually clipped to the 0 to 1 range. For 24 -bit RGB, 8 bits ( 256 levels) are used for each colour channel. This is done by multiplying $C_{S R G B}$ by 255 and rounding the result to an integer.

The clipping creates discontinuities that would result in sharp transition points in the colour spectrum. To smoothen the resulting colour spectrum, it is advisable to perform a 2D convolution of the data (e.g., conv2 function in Matlab or Octave). The final result is shown in Figure 11 and the GNU Octave script prepared to generate this graph is provided at the end of this technical note (1 page, blue background). Considering the limited gamut of the sRGB colour space, the spectral colour plot obtained with this method is globally satisfying. ${ }^{3}$

## 8. Summary

Colour spectrum representation was chosen as an example to illustrate the usefulness of the CIE 1931 chromaticity diagram. As you could expect, the use of the CIE 1931 XY colour space extends far beyond this simple case. To take but only few applications, colour LEDs data sheets are always provided with their bin specification in terms of xy coordinates in the CIE XY colour space; lightings are specified with a colour rendering index (CRI) calculated in the CIE colour space by comparing the colour rendering of a test source with an ideal source with a given correlated colour temperature (CCT); comparison of colour cameras passes through gamut analysis that is best displayed using the CIE chromaticity diagram.

[^2]

```
% Colour representation of the CIE }1931\mathrm{ chromaticity diagram
% Save the file as, e.g., 'CIE_1931_chromaticity_diagram.m'. Run this script in GNU Octave (v. 5.2.0).
% Prepared by: Christophe Yamahata, idonus Sàrl, June 2022 | Adapted from:
% https://ch.mathworks.com/matlabcentral/fileexchange/40640-computational-colour-science-using-matlab-2e
clear;
% Chromaticity coordinates of spectrum locus + line of purples
CIE_x = [ 0.17411 0.17380 0.17141 0.16888 0.16441 0.16111 0.15664 0.15099 0.14396 0.13550 0.12412 ...
0.10960 0.09129 0.06871 0.04539 0.04076 0.03620 0.03176 0.02749 0.02346 0.01970 0.01627 ...
0.01318 0.01048 0.00817 0.00628 0.00487 0.00398 0.00364 0.00386 0.00464 0.00601 0.00799 ...
0.01060 0.01387 0.01777 0.02224 0.02727 0.03282 0.03885 0.04533 0.05218 0.05932 0.06672 ...
0.07430 0.08205 0.08994 0.09794 0.10602 0.11416 0.12235 0.13055 0.13870 0.14677 0.15472 ...
0.16253 0.170240.17785 0.18539 0.19288 0.20031 0.207690.21503 0.22234 0.22962 0.23689 ...
0.24413 0.25136 0.25858 0.26578 0.27296 0.280130.287290.29445 0.30160 0.30876 0.31592 ...
0.32306 0.33021 0.33736 0.34451 0.35167 0.358810.365960.373100.38024 0.387380.39451 \ldots.
0.401630.40873 0.41583 0.42292 0.42999 0.437040.444060.451060.45804 0.464990.47190 \ldots.
0.47878 0.485610.49241 0.49915 0.50585 0.512490.519070.52560 0.53207 0.53846 0.54479 ...
0.551030.55719 0.563270.56926 0.57515 0.580940.586650.59222 0.59766 0.60293 0.60803 ...
0.612980.61778 0.62246 0.62704 0.64823 0.66576 0.680080.69151 0.70792 0.71903 0.72599 ...
0.733420.73469 0.67858 0.62248 0.56639 0.51030 0.454210.398110.34202 0.28593 0.22984 0.17411 ]';
CIE_y = [ 0.00496 0.00492 0.00510 0.00690 0.01086 0.01379 0.01771 0.02274 0.02970 0.03988 0.05780 ...
0.08684 0.13270 0.20072 0.29498 0.31698 0.33990 0.36360 0.387920.41270 0.43776 0.46295 ...
0.488210.51340 0.53842 0.56307 0.587120.610450.63301 0.654820.67590 0.69612 0.71534 ...
0.733410.75019 0.76561 0.779630.792110.802930.812020.81939 0.82516 0.82943 0.83227 ...
0.83380 0.83409 0.83329 0.83159 0.82918 0.826210.82277 0.81893 0.81478 0.81040 0.80586 \ldots..
0.80124 0.796520.79169 0.78673 0.78163 0.77640 0.77105 0.76559 0.76002 0.75433 0.74852 ...
0.742620.73661 0.73051 0.72432 0.71806 0.711720.705320.698840.692310.685710.67906 ...
0.672370.665630.65885 0.652030.645170.638290.63138 0.62445 0.61750 0.61054 0.60357 ...
0.596590.58961 0.582620.575630.56865 0.561670.554720.54777 0.54084 0.533930.52705 ...
0.52020 0.51339 0.50661 0.49989 0.49321 0.486590.480030.473530.467090.46073 0.45443 ...
0.448230.44210 0.43606 0.43010 0.42423 0.418460.412760.40719 0.40176 0.39650 0.39141 ...
0.38648 0.38171 0.37705 0.37249 0.35140 0.33401 0.319750.308340.292030.28094 0.27401 ...
0.26658 0.26531 0.23948 0.21363 0.18778 0.16192 0.13607 0.11022 0.08437 0.05852 0.03266 0.00496 ]';
```

\% CIE standard illuminant D65. For sRGB, the white point has chromaticity coordinates ( $0.3127,0.3290$ ).
\% This white point is also known as D65, which is an estimation of the white colour produced by mid-day sunlight.
D65 = [0.31271 0.32902];
\% Planckian locus: black body locus
planckian_locus_x = [ $0.73470 .72140 .70920 .69560 .68130 .66700 .65280 .63880 .62500 .6116 \ldots$
$0.59850 .58570 .57320 .56500 .56110 .54960 .52690 .50540 .48570 .46770 .45140 .43660 .4053 \ldots$
0.38050 .36050 .34390 .30830 .29970 .28540 .26510 .25130 .24130 .23380 .22790 .22320 .2146 ]';
planckian_locus_y = [ $0.26530 .27860 .29070 .30410 .31790 .33150 .34450 .35650 .36750 .3772 \ldots$
$0.38580 .39310 .39930 .40270 .40430 .40810 .41330 .41520 .41470 .41230 .40870 .40420 .3909 \ldots$
0.37670 .36330 .35070 .31860 .30960 .29380 .26900 .25060 .23650 .22530 .21630 .20890 .1950 ]';
\% sRGB gamut (colour triangle defined by the RGB primaries)
\% XYZ to sRGB transformation matrix
$\mathrm{M}=\left[\begin{array}{lll}3.2406255 & -1.537208 & -0.4986286 ;\end{array}\right.$
-0.9689307 1.8757561 0.0415175;
0.0557101 -0.2040211 1.0569959];
\%\%\% End of Part 1/3

```
% Gamut: colour triangle defined by the RGB primaries
Gamut =[inv(M)(:,1)/sum(inv(M)(:,1)) inv(M)(:,2)/sum(inv(M)(:,2)) inv(M)(:,3)/sum(inv(M)(:,3))];
gamut_x = [Gamut(1,:) Gamut(1,1)];
gamut_y = [Gamut(2,:) Gamut(2,1)];
function [rgb] = xy2srgb(xy, M)
    xyz = [xy(:,1) xy(:,2) 1-xy(:,1)-xy(:,2)];
    rgb = xyz*M';
    % Get brighter spectral colours, including a good yellow, scale up the linear RGB values,
    % allowing them to get higher than 1.0. Then, for each colour, scale all components back down,
    % if necessary, so that the maximum component value is 1.0.
    rgb = rgb / max(rgb(:));
    rgb = rgb*2.5;
    S = max(rgb,[],2);
    S = max(S,1);
    rgb = rgb ./ S;
    rgb = min(max(rgb,0),1);
    % Convert to nonlinear RGB values for the final result
    if(rgb <= 0.0031308)
    rgb = 12.92*rgb;
    else
    rgb = 1.055*rgb.^(1/2.4)-0.055;
    endif
end
N = length(CIE_x); i = 1; e = D65;
White_point = e;
sectors_mono = 8; % The larger these values, the better the graph quality
sectors_purple = 7; % (e.g., > 5 sectors gives good results).
sectors = sectors_mono;
xy4rgb = zeros(N*sectors*4,5,'double');
for w=1:N % equiv. wavelength
    w2 = mod(w,N)+1;
    a1 = atan2(CIE_y(w) -e(2),CIE_x(w) -e(1)); % start angle
    a2 = atan2(CIE_y(w2) -e(2),CIE_x(w2) -e(1)); % end angle
    r1 = ((CIE_x(w)-e(1))^2 + (CIE_y(w) -e(2))^2)^0.5; % start radius
    r2 = ((CIE_x(w2) - e(1))^2 + (CIE_y(w2) - e(2)}\mp@subsup{)}{}{\wedge}2\mp@subsup{)}{}{\wedge}0.5; % end radiu
    if(w >= N-10 || w <=10)
    sectors = sectors_purple;
    else
    sectors = sectors_mono;
    endif
for c=1:sectors % colourfulness (saturation)
    % patch polygon
    xy(1,1) = e(1)+r1*}\operatorname{cos}(a1)*c/sectors
    xy(1,2) = e(2)+r1*\operatorname{sin}(a1)*c/sectors;
    xy(2,1) =e(1)+r1*\operatorname{cos}(a1)*(c-1)/sectors;
```

\%\%\% End of Part 2/3

```
    xy(2,2) =e(2)+r1*\operatorname{sin}(a1)*(c-1)/sectors;
    xy(3,1) =e(1)+r2*}\operatorname{cos}(a2)*(c-1)/sectors
    xy(3,2) = e(2)+r2*}\operatorname{sin}(a2)*(c-1)/sectors
    xy(4,1) = e(1)+r2*}\operatorname{cos}(a2)*c/sectors
    xy(4,2) =e(2)+r2*}\operatorname{sin}(a2)*c/sectors
    % Compute RGB for vertices
    rgb = xy2srgb(xy, M);
    % Store the results
    xy4rgb(i:i+3,1:2) = xy(:,1:2);
    xy4rgb(i:i+3,3:5) = rgb;
    i = i + 4;
    end
end
```

\% Draw the CIE chromaticity diagram
min_x = 0; min_y = 0; max_x = 0.8; max_y = 0.9;
title_graph = 'CIE 1931 Chromaticity Diagram';
figure('Name', title_graph,'NumberTitle','off', 'ToolBar', 'none', 'Position', [600 50750 900]);
plot( [min_x; max_x], [0; 0], 'k-', 'LineWidth', 1.2, ...
[0; 0], [min_y; max_y], 'k-', 'LineWidth', 1.2, [0; 0.8], [1; 0.2], 'k--', 'LineWidth', 1);

```
% Will disappear progressively. We will plot it again at the end.
```

hold on
plot( White_point(1),White_point(2), 'o', 'MarkerFaceColor', 'black', 'MarkerSize',12, CIE_x, CIE_y,'k-', 'LineWidth', 3);
text(0.1,-0.17,'To interrupt this script, open the Command Window tab in $\{\backslash i$ itOctave $\}$ and type $\{\backslash \mathrm{bfCTRL}+\mathrm{C}\}$ ', 'Color',
[0.5 0.5 0.5],'FontSize',12);
set(gca,'FontSize',18); set(gca,'TickDir','out'); set(gca, 'XMinorTick', 'on', 'YMinorTick', 'on');
axis equal; axis([min_x max_x min_y max_y]);
xticks(0:0.1:max_x); yticks(0:0.1:max_y);
xlabel('x','FontSize',26); ylabel('y',''FontSize',26);
title(title_graph,'FontSize',28);
[rows cols] = size(xy4rgb);
$\mathrm{f}=\left[\begin{array}{ll}1 & 2 \\ 3 & 4\end{array}\right] ;$ v = zeros(4,3,'double');
for $\mathrm{i}=1: 4$ :rows
$v(:, 1: 2)=x y 4 r g b(i: i+3,1: 2)$;
patch('Vertices',v, 'Faces',f, 'EdgeColor','none', 'FaceVertexCData',xy4rgb(i:i+3,3:5),'FaceColor','interp');
refreshdata; drawnow;
end
\% Last layer: graphs on top of the coloured area
hold on
plot( planckian_locus_x, planckian_locus_y, 'Color', 0.6*[111],'LineWidth',2, ...
White_point(1),White_point(2), 'o', 'MarkerFaceColor', 'black', 'MarkerSize',12, ...
White_point(1),White_point(2), 'o', 'MarkerFaceColor', 'white', 'MarkerSize',8, ...
gamut_x, gamut_y, 'k-', 'LineWidth', 1, CIE_x, CIE_y,'k-', 'LineWidth', 3);
\%\%\% End of Part 3/3
\%\% This GNU Octave (v. 6.2.0) script was prepared by Christophe Yamahata, idonus Sàrl.
\%\% Copy the script and save the file as, e.g., 'RGB_to_spectrum.m'
clear;
lambda = [ 380:1:730]'; \% Wavelength range: $380 \mathrm{~nm}-730 \mathrm{~nm}$ (1 nm increment)
$R=[11122222333457891011131415171820212325262829313335373941434547495152 \ldots$ $5354555657585960616364666769707273757677777878797980808181797876757371706867 \ldots$ $655952463933262013700000000000000000000000000000000000000000000 \ldots$ 0000000000000000000000000005111622273238434954678093106119132145158171 ... 184191197204211218224231238244251251252252253253253254254255255255255255255255255 ... 255255255255255255255255255255255255255255255255255255255255255255255255255255255 ... 255255255255255255255255255255255255255255255255255251247243239235230226222218214 ... $209205200196191186182177173168164160157153149145141138134130127123120117114110107 \ldots$ 1041009794929986848178757370686563605855535048454442413938373534323130292827 ... 2624232221201918171615141312111010998877665 ]';
$\mathrm{G}=[00000000000000000000000000000000000000000000000000000000000000 \ldots$ $0000000000000000000816243240485664728084889296100103107111115119122124127 \ldots$ $129132135137140142145149152156159163167170174177181185189193197201205209213217221 \ldots$ 224228231235238241245248252255255255255255255255255255255255255255255255255255255 ... 255255255255255255255255255255255255255255255255255255255255255255255255255255255 ... 255255255255255255252248245242239235232229225222217212207202197192187182177172167 ... $1611561501451401341291231181101019384766859514234312724201714107300000000000 \ldots$ 0000000000000000000000000000000000000000000000000000000000000000 ... 00000000000000000000000000000000000 ]';
B = $[34567889101112151720232628313436394245495255586165687175798286909498101 \ldots$ $105109111114116118121123125127130132136140144148152155159163167171175178182185189 \ldots$ 193196200203207212217221226231236241245250255253250248245243241238236233231227222 ... 218213209205200196191187185183180178176174172169167165165165165165166166166166166 ... $166167168169170172173174175176177178178179180181181182183183184184183183183183182 \ldots$ $18218218118117817517216916616215915615315013512010590756045301500000000000000 \ldots$ 00000000000000000000000000000000000000000000000715222937445158667373 ... $7474747575757576767575747373727170706967666463616058575554535150494846454442 \ldots$ $41403937363534333130292827262524232221201918171615141312111099887766554443 \ldots$ 33332222222111111111111111 ]';
$R G B=[R G B] ;$
\% First, perform a convolution to smoothen the data
RGB = conv2(RGB, ones(31,1)/31,'same');

## \% Draw the graph

figure('NumberTitle','off', 'Name', "Rendering of the solar spectrum with consideration of the sRGB gamut limitations", 'ToolBar', 'none', 'MenuBar', 'none', 'Position', [100 2001000 800]);
plot(lambda(:),RGB(:,1), 'r', lambda(:),RGB(:,2), 'g',lambda(:),RGB(:,3), 'b');
set(gca,'FontSize',16,'TickDir','out', 'XMinorTick', 'on', 'YMinorTick', 'on');
title (\{'RGB values \rm\{calculated from the\}';
' $\mathrm{Irm}\{\backslash i t\{C I E 1931$ chromaticity diagram $\}\} \backslash r m\{$ and D65 standard illuminant\}'\});
legend('R', 'G','B', 'location', 'northwest', 'box','off');
axis([380 7300255$])$; xticks(400:50:720); yticks(0:50:250); xlabel('Wavelength (nm)','FontSize',22);

## \% Draw the spectrum

intensity_R=RGB(:,1)/255; intensity_G=RGB(:,2)/255; intensity_B=RGB(:,3)/255;
colormap ([intensity_R, intensity_G,intensity_B]); caxis([380 730]);
h = colorbar ("southoutside"); set(h,'fontsize',16);

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[^0]:    ${ }^{1}$ In CIE's convention, lower case letters represent the chromaticity coordinates in a colour space of the same name written in upper case letters (e.g., $r, g$, and $b$ coordinates in the CIE RGB colour space).

[^1]:    ${ }^{2}$ GNU Octave, scientific programming language. Version used: GNU Octave 6.2.0.
    [ https://octave.org/]

[^2]:    ${ }^{3}$ Useful reading in relation with section 7: "Making Color Spectrum Plots - Part 2," Steve Eddins, July 20, 2020.
    [ https://blogs.mathworks.com/steve/2020/07/20/making-color-spectrum-plots-part-2/ ]

